Fundamental Symmetries through the lens of a neutron

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Planetary orbits

- Kepler's first law of planetary motion: the orbit is an ellipse with the sun at one focus
- Symmetry: the long axis can point in any direction
- Broken symmetry: at any given time, the long axis points in a specific direction
- Sensitive to perturbations...



Planetary orbits

- Precession of Mercury's orbit
 - 565" per earth century observed
 - Outer planets explained 527" per earth century
- New particle? Planet Vulcan?
- New model? General relativity
- Use symmetry to find new physics!



Standard Model of Particle Physics



- Many questions: gravity? dark matter? missing antimatter? finetuning? 26 free parameters?
- We haven't seen new particles at the LHC (yet)...
- Is there an underlying framework?
- Is there an underlying symmetry?

Symmetries in nature

- Spacetime symmetries
 - Time, Space, Rotations
- Permutation symmetries
 - Fermi-Dirac/Bose-Einstein statistics
- Internal symmetries
 - Charge, Lepton #, Baryon #, ...
- Discrete symmetries
 - Charge, Parity, Time



Outline

- P-violation and the weak interaction
- CP-violation and the electric dipole moment
- B-violation and oscillations

Neutron beta decay

- Weak interaction mixes quarks
 - "beta" = electron/positron emitted to conserve charge, along with (anti-)neutrino
- Neutron is simplest "nucleus" to beta decay
- CKM Unitarity test
 - If you rotate, and unrotate, you should get back where you started
 - As you rotate, your yardstick shouldn't change length



Failure means some new physics is missing...

The electroweak interaction

• General interaction vertex

Scalar	$\overline{\psi}\phi$
Pseudoscalar	$ar{\psi}\phi$
Vector	$ar{\psi}\gamma^\mu\phi$
Axial Vector	$ar{\psi}\gamma^{\mu}\gamma^{5}\phi$
Tensor	$\bar{\psi}(\gamma^{\mu}\gamma^{\nu}-\gamma^{\nu}\gamma^{\mu})\phi$

• Could other currents be participating?

• Weak interaction has experimentally observed form:

$$J_{had}^{\mu} = V_{ud}\bar{u}(\gamma^{\mu} - \gamma^{\mu}\gamma^{5})d$$
$$J_{lep}^{\mu} = \bar{e}(\gamma^{\mu} - \gamma^{\mu}\gamma^{5})v_{e}$$

- Helicity operator $1 \gamma^5$ projects out left-handed particles/right-handed antiparticles
- Weak interaction V-A is maximally parity-violating

Nuclear beta decay



Unitarity status

- CKM unitarity appeared valid until 2018
- Reanalysis of Δ_R^V : improved uncertainty and dramatically shifted value of V_{ud} from 0⁺ \rightarrow 0⁺ superallowed decays
 - 3σ violation!
- Strong motivation for new complementary measurements



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Measurables in neutron decay

- Angular correlations $dW \propto 1 + a \frac{\vec{p}_e \cdot \vec{p}_v}{E_e E_u} + b \frac{m_e}{E_e} + \langle \vec{\sigma}_n \rangle \cdot \left(A \frac{\vec{p}_e}{E_e} + B \frac{\vec{p}_v}{E_u} + D \frac{\vec{p}_e \times \vec{p}_v}{E_e E_u} \right) \qquad \text{neutron}$ electron • A, a sensitive to A, V interactions В • $A = -2 \frac{\lambda^2 + \lambda}{1 + 3\lambda^2}$ $a = \frac{1 - \lambda^2}{1 + 3\lambda^2}$ $\lambda = \frac{g_A}{g_V}$ • **B**, **b** sensitive to **S**, **T** • Decay lifetime antineutrino $\boldsymbol{\tau}^{-1} = \boldsymbol{W} \propto (\boldsymbol{V}_{\boldsymbol{ud}})^2 (1 + 3(\boldsymbol{\lambda})^2)$ • 2 unknowns = 2 observables proto
- Improved LQCD calcs of g_A , g_S , g_T

Chang et al, Nature 558 (2018) 91-94 Gupta et al, PRD 98 (2018) 034503

Vud from neutron decay



Sensitivity to Scalars and Tensors currents

• Asymmetries:
$$\alpha_{\text{meas}}(E_e) = \frac{\alpha(E_e)}{1+bm_e/E_e}$$

• Spectral measurement: $1 + b m_e/E_e$ distortion to decay rate

• **B** (b_v), **b** linear sensitivity to BSM **S**,**T**:
•
$$b^{BSM} = \frac{2}{1+3\lambda^2} [g_S \epsilon_S - 12\lambda g_T \epsilon_T]$$

• $b^{BSM}_{\nu} = \frac{2}{1+3\lambda^2} [\lambda g_S \epsilon_S - 4g_T \epsilon_T (1+2\lambda)]$



Gupta et al, PRD 98 (2018) 034503

The neutron

- Source: Freed from atoms
- Mass: 1.0087 a.m.u. (Proton: 1.0073 a.m.u)
- Spin: $\frac{1}{2}$
- Gravity: 100 neV per m
- Electromagnetism: Electric charge: 0 Magnetic dipole moment: 60 neV per 1 T Electric dipole moment: 0 e-cm (?)
- Weak: beta decay lifetime of about 15 minutes
- Strong: neutrons can interact with matter



Making free neutrons

• Lots of energy required (MeV scale) to free neutrons from atom



- Experiments need slow neutrons
 - Longer wavelengths = easier to guide
 - Slower = more decays in your experiment

Class	Energy	Source
Fast	> 1 MeV	Fission / spallation
Slow	eV – keV	Moderation
Thermal	0.025 eV	Thermal equilibrium
Cold	µeV – meV	Cold moderation
Ultracold	< 300 neV	Superthermal
		process

Neutrons and matter

- Coherent scattering from many nuclei $\psi_n(r) = e^{ikz} + b \frac{e^{ikr}}{r}$
- Neutron wavefunction satisfies wave equation $(\nabla^2 + k^2)\psi_n(r) = 4 \pi N b(r) \psi_n(r)$
- Snell's Law = neutron guides $n = \sqrt{1 \frac{4\pi Nb}{k^2}}$ $n_1 \sin \theta_1 = n_2 \sin \theta_2$ $\theta_c = \lambda \sqrt{Nb/\pi}$ $\lambda_c = \sqrt{\pi/Nb} > 500 \text{ Å}$ $\theta < \theta_c$
- Ultracold neutrons have $\theta_c = 90^{\circ}$

θ

 $\theta > \theta_c$



Neutron Lifetime

- "Beam" technique: count the dying
- "Bottle" technique: count the survivors
- 8.6 s (~4 σ) discrepancy between methods!



Neutron beta decay

• Understanding this process is critical to many other applications:



Beam Lifetime Experiments: BL2



Slide adapted from S. Hoogerheide, PPNS2018 A. Yue et al, Phys. Rev. Lett. **111** 222501 (2013)

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Bottle lifetime experiments

• Measure neutrons remaining after storage: $N(\Lambda t) = N_{c} \rho^{-\Delta t/\tau_{s}}$

$$(\Delta t) = N_0 e^{-\Delta t_{\beta}}$$
$$\frac{1}{\tau_s} = \frac{1}{\tau_{\beta}} + \frac{1}{\tau_{other}}$$

- Loss mechanisms: upscattering on walls, capture on walls, gaps in walls, quasibound escape, depolarization, residual gas...
- Wall losses depend on velocity, collision rate, loss factor:

$$\tau_i^{-1} = \gamma(\nu)\mu_i(\nu)$$





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Magnetic Bottle: UCN₇



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"Fundamental Symmetries" Exotic Beam Summer School, ORNL, July 24-29

The β Asymmetry: UCNA

- $W \propto 1 + \frac{v}{c} \langle P \rangle A(E) \cos \theta$
- Magnetic spectrometer: $\langle \cos \theta \rangle = \pm \frac{1}{2}$
- Measure asymmetry: 2 detectors, 2 spin directions

$$A_{exp} = \frac{N^{+} - N^{-}}{N^{+} + N^{-}} = \frac{1}{2} \frac{v}{c} PA$$

Scintillator • Cancel systematics with Super-Ratio $R(E) = \frac{N(E)_1^+ N(E)_2^-}{N(E)_2^+ N(E)_2^+}$

$$A_{SR} = \frac{1 - \sqrt{R}}{1 + \sqrt{R}} = \frac{v}{c} \langle P \rangle A(E) \cos \theta$$

• "Super-sum" removes distortion from A

$$C = \frac{1}{2}\sqrt{N(E)_{1}^{+}N(E)_{2}^{-}} + \frac{1}{2}\sqrt{N(E)_{1}^{-}N(E)_{2}^{+}}$$

Brown et al, PRC 97 035505 (2018) Hickerson et al, PRC 96 (2017) 042501



Array

MWPC



Maerkisch et al, Phyr. Rev. Lett. 122 242501 (2019)

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The β - ν Correlation: aCORN



Slides adapted from F. E. Wietfeldt G. Darius et al, Phys. Rev. Lett. 119 042502 (2017)

The β - ν Correlation: Nab at the SNS



Bowman, J Res NIST 110 40 (2005) Pocanic et al, NIMA 611 211 (2009) Baessler et al, J Phys G 41 114003 (2014)

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Nab at the SNS

- 1. Unpolarized neutrons enter experiment region
- 2. Neutrons decay, charged particles are trapped by magnetic fields
- 3. Electrons and protons are guided along magnetic field lines to detectors
- 4. Detectors determine electron energy and proton time of flight



Nab detectors

- Challenge to detect both: protons ~ 0.8 keV (requires accelerating voltage of -30 kV) electrons ~ 800 keV
- Demonstrated with UCN!
- Fully instrumented system acceptance testing nearly complete at LANL
- Precision characterizations next: deadlayer, charge collection profile, calibration, cross-talk...

Salas-Bacci et al, NIMA 735 (2014) 408; Sjue et al, RSI 86 (2015) 023102; LJB et al, NIMA 849 (2017) 83; LJB et al, J. Phys: Conf Ser. 876 012005 (2017) LJB et al, Hyp. Int. 240 1 (2019)



Nab magnetic spectrometer

- Charged particles spiral along B field lines: $qv_{\perp}B = \frac{mv_{\perp}^2}{r}$
- Magnetic moment is an adiabatic invariant: $\mu = \frac{mv_{1,\perp}^2/2}{B_1} = \frac{mv_{2,\perp}^2/2}{B_2}$
- Conservation of energy: transform v_{\perp} to v_{\parallel}
- Increase B: decrease v_{\parallel} (particle reflects)
- Decrease B: motion becomes parallel to field...



Nab magnetic spectrometer

- 4 T magnetic mirror (filter): reflect/ignore protons with small v_{\parallel}
- 0.2 T TOF region: proton motion nearly parallel to B = spectrometer axis
- Relate time of flight (measured) to proton momentum (desired):

$$t_p = L \frac{m_p}{p_p} = \frac{f(\cos \theta)}{p_p}$$



Nab spectrometer now commissioning

- Performance success—excellent agreement with calculations
- Precision mapping to be completed *imminently*





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The early universe

- Big Bang Nucleosynthesis calculations: how light elements were formed
- Cosmic Microwave Background: "Big Bang's echo"
- Where's the antimatter?
 - Cosmic rays?
 - Antimatter galaxies?
 - BBN, CMB agree:

$$\eta\equivrac{n_B-n_{\overline{B}}}{s}$$
 ~ 10⁻¹⁰





Look for clues in symmetries

- Internal symmetry: **Baryon number B**
 - Quarks: $\mathbf{B} = +\frac{1}{3}$
 - Antiquarks: $\mathbf{B} = -\frac{1}{3}$
- Big Bang should have produced matter = antimatter $\rightarrow \mathbf{B} = \mathbf{0}$
- First condition: we must have a process that does not conserve **B** to create an excess
 - $X + Y \rightarrow \mathbf{B}$
 - Has not been observed...



Andrei Sakharov

Look for clues in symmetries

- Second Condition: Production of matter must be favored over antimatter
- C-symmetry must be broken $X + Y \rightarrow \mathbf{B} > \overline{X} \rightarrow \overline{Y} + \overline{B}$
- CP-symmetry must be broken

$$(CP - X \rightarrow B_L + X \rightarrow B_R > Conjugates) \quad \overline{X} \rightarrow B_L + \overline{X} \rightarrow \overline{B}_R$$

• **Third Condition:** Excess **B** shouldn't be washed out – need this process to occur out of thermal equilibrium



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Electric dipole moments and CP violation

• Non-zero electric dipole moment in fundamental particle: $H = -\frac{1}{2} \left(\vec{d} \cdot \vec{E} + \vec{\mu} \cdot \vec{B} \right)$ \vec{d} \vec{d} • P-violation, T-violation and therefore **CP**-violation Ù

 $\vec{\mu}$

Electric dipole moments



Chupp, Fierlinger, Ramsey-Musolf, and Singh, Rev. Mod. Phys. 91 1 (2019)

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Neutron electric dipole moment

- Neutron has magnetic dipole moment: feels torque perpendicular to magnetic field
- Spin precesses with frequency $h\nu = 2\mu_n B$
- If neutron had an EDM, spin precesses around electric field:

$h\nu = 2d_n E$

• Tactic: compare frequency with electric/ magnetic fields parallel/antiparallel

 $h\nu_{\pm} = 2\mu_n B \pm 2d_n E$ $\Delta\nu = \frac{4d_n E}{h}$

$$\Delta v = 7.5 \text{ nHz} \text{ for } d_n = 10^{-28} \text{ e-cm}, E = 75 \text{ kV/cm}$$



Limit on the neutron EDM



- First measured by Smith, Purcell and Ramsey at ORNL
 - Studying parity violation in neutron scattering
- Neutron EDM extremely small: "Strong CP problem" in QCD Langrangian
- Motivates "axions" (dark matter candidate)

Smith, Purcell, and Ramsey, Phys Rev 108 120 (1957)

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Ramsey's Method of Separated Oscillating Fields

- 1. Prepare "spin up" neutron
- 2. Apply 90° spin flip
- 3. Spin freely precesses by $\omega \Delta t$
- 4. Apply 90° spin flip (to "spin down")
- 5. Measure "spin up" vs "spin down"



Т

Т

Present best nEDM limit

- Trapped 0.5 UCN/cc, stored for ~100 s
- $E \sim 10 \text{ kV/cm}$
- Result: $d_n = -0.21 \pm 1.82 \times 10^{-26} \text{ e-cm}$

UCN from superfluid He (He-II)

- UCN production via "superthermal process"
- 8.9 Å (1 meV/11 K) cold neutron emits phonon in He-II, becomes UCN (500 Å/300 neV/3.5 mK)
- He-II cooled to ~1 K: Reverse process is suppressed by Boltzmann factor e^{-E/k_BT}

nEDM at the SNS

- Use helium to improve statistical precision
 - E: LHe = high voltage insulator, >85 kV/cm achieved
 - N: Higher densities using He-II
 - t: temp < 0.5 K allows storage times of ~ 1000 s

- and reduce systematic uncertainty
 - "Geometric phase": interaction of $\vec{v} \times \vec{E} = \vec{B}_E$ and magnetic gradients
 - Limit leakage currents from electric field (which may produce magnetic fields)
 - Multiple measurement cells with opposite $\vec{E} \cdot \vec{B}$
 - Multiple measurement techniques

Golub and Lamoreaux, Phys Rep 237 1 (1994) Ito et al, Rev Sci Instrum 87 045113 (2016)

SNS nEDM

- He-II is UCN source
- ³He captures neutrons with strong spin-dependence
- He-II acts as a scintillator
- ³He (95% polarized) used as a co-magnetometer

Extracting the nEDM

- ³He capture is strongly spindependent
 - 69 kb if spins anti-aligned
 - 11 b if spins aligned
- Time evolution of angle:

$$\theta_{n3} = |\gamma_n - \gamma_3| B_0 t \pm \frac{ed_n |E|}{\hbar} t$$

- $\cos(\omega_n \omega_3)$ variation in reaction rate
- Beat frequency: 9 Hz @ 30 mG
- ³He as co-magnetometer
 - ³He shielded by atomic electrons
 - ³He precession measured directly by SQUID pickup

nEDM now being constructed

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Neutron anti-neutron oscillations

$$\widehat{H} = \begin{pmatrix} m + \vec{\mu}(\vec{B} \cdot \vec{\sigma}) & \varepsilon \\ \varepsilon & m - \vec{\mu}(\vec{B} \cdot \vec{\sigma}) \end{pmatrix}$$

$$\Delta E = E_n - E_{\bar{n}} = 2\mu B$$

$$\Delta E (1 nT) \sim 10^{-22} \text{ MeV} \qquad \varepsilon < 10^{-29} \text{ MeV}$$

• Need "Quasi-free limit": Uncertainty principle = $\Delta E \Delta t \ll \hbar$

2

• ILL: neutron TOF ~ 0.1 s, B~10 nT

• In this limit
$$P_{n \to \overline{n}}(t) = \left(\frac{t_{free}}{\tau_{n \to \overline{n}}}\right)$$

Free $n \rightarrow \overline{n}$ Search at ILL

Improving limits on $n \rightarrow \overline{n}$: European Spallation Source

- Similar time-averaged brightness to ILL
 - Pulsed = bkgd rejection
- >200 m beamguides possible
- Moderator, neutron optics, detector design study
 - ~1000x ILL possible

Moderator configuration	Sensitivity Nt ² in ILL units/yr
ESS TDR 2013	250
Option of large LD ₂ source	550
Flat "pancake" with h = 3 cm	200
ANNI with BF1 Source	0.8

Adapted from M. Frost, INT 17-69W

\bar{n} annihilation detection

- Maintain "background free", improve efficiency
- Example "pion star" $\overline{n} + {}^{12}C \rightarrow {}^{11}C$ $+3\pi^0 + \pi^+ + \pi^-$
- Goal background suppression $< 10^{-8}$ Hz
- Key backgrounds:
 - Gamma backgrounds
 - Neutron capture
 - High energy n,p
 - Beta delayed n
 - Cosmic rays
 - Dominant, scales with size

Detector study by A.R. Young, R.W. Pattie Jr, D.G. Phillips II

$n \rightarrow \overline{n}$ in Nuclei

• Best limits from SuperKamiokande: $\tau_{n \to \bar{n}} > 2.7 \times 10^8$ s

24 candidates, 24.1 expected backgrounds

Adapted from J. Barrow, BLV 2017 Abe et al, PRD 91 072006 (2015)

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The Deep Underground Neutrino Experiment and Proposed $n - \bar{n}$ Search

Babu, Dev, Fortes, and Mohapatra- DOI:10.1103/PhysRevD.87.115019

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Mirror neutron oscillations

- "Mirror matter": hidden sector dark matter candidate
- Multiple Baryon Numbers, with overall conservation of $B_{tot} = B + B'$
- Experimental signature: oscillation into invisible "mirror neutrons"
 - All previous searches performed in ultracold neutron bottle experiments
- Various mechanisms to induce oscillations, using magnetic fields
 - some suggested to explain the neutron lifetime discrepancy
 - can consider mixing with antineutrons

$$\mathcal{H}_{int} = \begin{pmatrix} m + \mu \boldsymbol{\sigma} \cdot \boldsymbol{B} & \eta \boldsymbol{\sigma} \cdot [\boldsymbol{B} \pm \boldsymbol{B}'] + \varepsilon \\ \eta \boldsymbol{\sigma} \cdot [\boldsymbol{B} \pm \boldsymbol{B}'] + \varepsilon & m' + \mu' \boldsymbol{\sigma} \cdot \boldsymbol{B}' \end{pmatrix}$$

Cold $n \rightarrow n'$ searches

neutrons transform into mirror neutrons

- Add B field control P(n and high suppression "wall" to GP-SANS
- First small scale experiment this summer at SNS

Summary

- Tremendous opportunity to search for influence of new physics using low energy precision tests of fundamental symmetries
- Violation of CKM unitarity by 3 σ demands new precision measurements using independent approaches
 - New precision neutron measurements on the horizon including Nab at SNS
- Matter asymmetry in universe remains one of the most important questions in science
 - Electric dipole moment searches such as nEDM at SNS will probe interesting parameter space for theories predicting CP-violation
 - Observation of neutron oscillations would demonstrate baryon number violating process
- Lots of exciting developments in neutrons and nuclei.... stay tuned!