## Fundamental Symmetries through the lens of a neutron

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Exotic Beams Summer School June 24, 2019

## Planetary orbits

- Kepler's first law of planetary motion: the orbit is an ellipse with the sun at one focus
- Symmetry: the long axis can point in any direction
- Broken symmetry: at any given time, the long axis points in a specific direction
- Sensitive to perturbations...



## Planetary orbits

- Precession of Mercury's orbit
  - 565" per earth century observed
  - Outer planets explained 527" per earth century
- New particle? Planet Vulcan?
- New model? General relativity
- Use symmetry to find new physics!



## Standard Model of Particle Physics



- Many questions: gravity? dark matter? missing antimatter? finetuning? 26 free parameters?
- We haven't seen new particles at the LHC (yet)...
- Is there an underlying framework?
- Is there an underlying symmetry?

## Symmetries in nature

- Spacetime symmetries
  - Time, Space, Rotations
- Permutation symmetries
  - Fermi-Dirac/Bose-Einstein statistics
- Internal symmetries
  - Charge, Lepton #, Baryon #, ...
- Discrete symmetries
  - Charge, Parity, Time



#### Outline

- P-violation and the weak interaction
- CP-violation and the electric dipole moment
- B-violation and oscillations

## Neutron beta decay

- Weak interaction mixes quarks
  - "beta" = electron/positron emitted to conserve charge, along with (anti-)neutrino
- Neutron is simplest "nucleus" to beta decay
- CKM Unitarity test
  - If you rotate, and unrotate, you should get back where you started
  - As you rotate, your yardstick shouldn't change length



Failure means some new physics is missing...

#### The electroweak interaction

• General interaction vertex

Scalar	$\overline{\psi}\phi$
Pseudoscalar	$ar{\psi}\phi$
Vector	$ar{\psi}\gamma^\mu\phi$
Axial Vector	$ar{\psi}\gamma^{\mu}\gamma^{5}\phi$
Tensor	$\bar{\psi}(\gamma^{\mu}\gamma^{\nu}-\gamma^{\nu}\gamma^{\mu})\phi$

• Could other currents be participating?

• Weak interaction has experimentally observed form:

$$J_{had}^{\mu} = V_{ud}\bar{u}(\gamma^{\mu} - \gamma^{\mu}\gamma^{5})d$$
$$J_{lep}^{\mu} = \bar{e}(\gamma^{\mu} - \gamma^{\mu}\gamma^{5})v_{e}$$

- Helicity operator  $1 \gamma^5$  projects out left-handed particles/right-handed antiparticles
- Weak interaction V-A is maximally parity-violating

## Nuclear beta decay



## Unitarity status

- CKM unitarity appeared valid until 2018
- Reanalysis of  $\Delta_R^V$ : improved uncertainty and dramatically shifted value of  $V_{ud}$  from 0<sup>+</sup>  $\rightarrow$  0<sup>+</sup> superallowed decays
  - $3 \sigma$  violation!
- Strong motivation for new complementary measurements



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#### Measurables in neutron decay

- Angular correlations  $dW \propto 1 + a \frac{\vec{p}_e \cdot \vec{p}_v}{E_e E_u} + b \frac{m_e}{E_e} + \langle \vec{\sigma}_n \rangle \cdot \left( A \frac{\vec{p}_e}{E_e} + B \frac{\vec{p}_v}{E_u} + D \frac{\vec{p}_e \times \vec{p}_v}{E_e E_u} \right) \qquad \text{neutron}$ electron • A, a sensitive to A, V interactions В •  $A = -2 \frac{\lambda^2 + \lambda}{1 + 3\lambda^2}$   $a = \frac{1 - \lambda^2}{1 + 3\lambda^2}$   $\lambda = \frac{g_A}{g_V}$ • **B**, **b** sensitive to **S**, **T** • Decay lifetime antineutrino  $\boldsymbol{\tau}^{-1} = \boldsymbol{W} \propto (\boldsymbol{V}_{\boldsymbol{ud}})^2 (1 + 3(\boldsymbol{\lambda})^2)$ • 2 unknowns = 2 observables proto
- Improved LQCD calcs of  $g_A$ ,  $g_S$ ,  $g_T$

Chang et al, Nature 558 (2018) 91-94 Gupta et al, PRD 98 (2018) 034503

## Vud from neutron decay



#### Sensitivity to Scalars and Tensors currents

• Asymmetries: 
$$\alpha_{\text{meas}}(E_e) = \frac{\alpha(E_e)}{1+bm_e/E_e}$$

• Spectral measurement:  $1 + b m_e/E_e$ distortion to decay rate

• **B** (b<sub>v</sub>), **b** linear sensitivity to BSM **S**,**T**:  
• 
$$b^{BSM} = \frac{2}{1+3\lambda^2} [g_S \epsilon_S - 12\lambda g_T \epsilon_T]$$
  
•  $b^{BSM}_{\nu} = \frac{2}{1+3\lambda^2} [\lambda g_S \epsilon_S - 4g_T \epsilon_T (1+2\lambda)]$ 



Gupta et al, PRD 98 (2018) 034503

#### The neutron

- Source: Freed from atoms
- Mass: 1.0087 a.m.u. (Proton: 1.0073 a.m.u)
- Spin:  $\frac{1}{2}$
- Gravity: 100 neV per m
- Electromagnetism: Electric charge: 0 Magnetic dipole moment: 60 neV per 1 T Electric dipole moment: 0 e-cm (?)
- Weak: beta decay lifetime of about 15 minutes
- Strong: neutrons can interact with matter



## Making free neutrons

• Lots of energy required (MeV scale) to free neutrons from atom



- Experiments need slow neutrons
  - Longer wavelengths = easier to guide
  - Slower = more decays in your experiment

Class	Energy	Source
Fast	> 1 MeV	Fission / spallation
Slow	eV – keV	Moderation
Thermal	0.025 eV	Thermal equilibrium
Cold	µeV – meV	Cold moderation
Ultracold	< 300 neV	Superthermal
		process

#### Neutrons and matter

- Coherent scattering from many nuclei  $\psi_n(r) = e^{ikz} + b \frac{e^{ikr}}{r}$
- Neutron wavefunction satisfies wave equation  $(\nabla^2 + k^2)\psi_n(r) = 4 \pi N b(r) \psi_n(r)$
- Snell's Law = neutron guides  $n = \sqrt{1 \frac{4\pi Nb}{k^2}}$   $n_1 \sin \theta_1 = n_2 \sin \theta_2$   $\theta_c = \lambda \sqrt{Nb/\pi}$   $\lambda_c = \sqrt{\pi/Nb} > 500 \text{ Å}$  $\theta < \theta_c$
- Ultracold neutrons have  $\theta_c = 90^{\circ}$

θ

 $\theta > \theta_c$ 



#### Neutron Lifetime

- "Beam" technique: count the dying
- "Bottle" technique: count the survivors
- 8.6 s (~4 σ) discrepancy between methods!



## Neutron beta decay

• Understanding this process is critical to many other applications:



## Beam Lifetime Experiments: BL2

![](_page_19_Figure_1.jpeg)

Slide adapted from S. Hoogerheide, PPNS2018 A. Yue et al, Phys. Rev. Lett. **111** 222501 (2013)

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## Bottle lifetime experiments

• Measure neutrons remaining after storage:  $N(\Lambda t) = N_{c} \rho^{-\Delta t/\tau_{s}}$ 

$$(\Delta t) = N_0 e^{-\Delta t_{\beta}}$$
$$\frac{1}{\tau_s} = \frac{1}{\tau_{\beta}} + \frac{1}{\tau_{other}}$$

- Loss mechanisms: upscattering on walls, capture on walls, gaps in walls, quasibound escape, depolarization, residual gas...
- Wall losses depend on velocity, collision rate, loss factor:

$$\tau_i^{-1} = \gamma(\nu)\mu_i(\nu)$$

![](_page_20_Figure_6.jpeg)

![](_page_21_Picture_0.jpeg)

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Magnetic Bottle: UCN<sub>7</sub>

![](_page_22_Figure_1.jpeg)

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"Fundamental Symmetries" Exotic Beam Summer School, ORNL, July 24-29

## The β Asymmetry: UCNA

- $W \propto 1 + \frac{v}{c} \langle P \rangle A(E) \cos \theta$
- Magnetic spectrometer:  $\langle \cos \theta \rangle = \pm \frac{1}{2}$
- Measure asymmetry: 2 detectors, 2 spin directions

$$A_{exp} = \frac{N^{+} - N^{-}}{N^{+} + N^{-}} = \frac{1}{2} \frac{v}{c} PA$$

Scintillator • Cancel systematics with Super-Ratio  $R(E) = \frac{N(E)_1^+ N(E)_2^-}{N(E)_2^+ N(E)_2^+}$ 

$$A_{SR} = \frac{1 - \sqrt{R}}{1 + \sqrt{R}} = \frac{v}{c} \langle P \rangle A(E) \cos \theta$$

• "Super-sum" removes distortion from A

$$C = \frac{1}{2}\sqrt{N(E)_{1}^{+}N(E)_{2}^{-}} + \frac{1}{2}\sqrt{N(E)_{1}^{-}N(E)_{2}^{+}}$$

Brown et al, PRC 97 035505 (2018) Hickerson et al, PRC 96 (2017) 042501

![](_page_23_Figure_10.jpeg)

Array

MWPC

![](_page_24_Figure_0.jpeg)

Maerkisch et al, Phyr. Rev. Lett. 122 242501 (2019)

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## The $\beta$ - $\nu$ Correlation: aCORN

![](_page_25_Figure_1.jpeg)

Slides adapted from F. E. Wietfeldt G. Darius et al, Phys. Rev. Lett. 119 042502 (2017)

#### The $\beta$ - $\nu$ Correlation: Nab at the SNS

![](_page_26_Figure_1.jpeg)

Bowman, J Res NIST 110 40 (2005) Pocanic et al, NIMA 611 211 (2009) Baessler et al, J Phys G 41 114003 (2014)

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## Nab at the SNS

- 1. Unpolarized neutrons enter experiment region
- 2. Neutrons decay, charged particles are trapped by magnetic fields
- 3. Electrons and protons are guided along magnetic field lines to detectors
- 4. Detectors determine electron energy and proton time of flight

![](_page_27_Figure_5.jpeg)

## Nab detectors

- Challenge to detect both: protons ~ 0.8 keV (requires accelerating voltage of -30 kV) electrons ~ 800 keV
- Demonstrated with UCN!
- Fully instrumented system acceptance testing nearly complete at LANL
- Precision characterizations next: deadlayer, charge collection profile, calibration, cross-talk...

Salas-Bacci et al, NIMA 735 (2014) 408; Sjue et al, RSI 86 (2015) 023102; LJB et al, NIMA 849 (2017) 83; LJB et al, J. Phys: Conf Ser. 876 012005 (2017) LJB et al, Hyp. Int. 240 1 (2019)

![](_page_28_Picture_6.jpeg)

## Nab magnetic spectrometer

- Charged particles spiral along B field lines:  $qv_{\perp}B = \frac{mv_{\perp}^2}{r}$
- Magnetic moment is an adiabatic invariant:  $\mu = \frac{mv_{1,\perp}^2/2}{B_1} = \frac{mv_{2,\perp}^2/2}{B_2}$
- Conservation of energy: transform  $v_{\perp}$  to  $v_{\parallel}$
- Increase B: decrease  $v_{\parallel}$  (particle reflects)
- Decrease B: motion becomes parallel to field...

![](_page_29_Picture_6.jpeg)

## Nab magnetic spectrometer

- 4 T magnetic mirror (filter): reflect/ignore protons with small  $v_{\parallel}$
- 0.2 T TOF region: proton motion nearly parallel to B = spectrometer axis
- Relate time of flight (measured) to proton momentum (desired):

$$t_p = L \frac{m_p}{p_p} = \frac{f(\cos \theta)}{p_p}$$

![](_page_30_Figure_5.jpeg)

## Nab spectrometer now commissioning

- Performance success—excellent agreement with calculations
- Precision mapping to be completed *imminently*

![](_page_31_Figure_3.jpeg)

![](_page_31_Picture_4.jpeg)

#### Outline

- P-violation and the weak interaction
- CP-violation and the electric dipole moment
- B-violation and oscillations

## The early universe

- Big Bang Nucleosynthesis calculations: how light elements were formed
- Cosmic Microwave Background: "Big Bang's echo"
- Where's the antimatter?
  - Cosmic rays?
  - Antimatter galaxies?
  - BBN, CMB agree:

$$\eta\equivrac{n_B-n_{\overline{B}}}{s}$$
 ~ 10<sup>-10</sup>

![](_page_33_Figure_8.jpeg)

![](_page_33_Figure_9.jpeg)

## Look for clues in symmetries

- Internal symmetry: **Baryon number B** 
  - Quarks:  $\mathbf{B} = +\frac{1}{3}$
  - Antiquarks:  $\mathbf{B} = -\frac{1}{3}$
- Big Bang should have produced matter = antimatter  $\rightarrow \mathbf{B} = \mathbf{0}$
- First condition: we must have a process that does not conserve **B** to create an excess
  - $X + Y \rightarrow \mathbf{B}$
  - Has not been observed...

![](_page_34_Picture_8.jpeg)

Andrei Sakharov

## Look for clues in symmetries

- Second Condition: Production of matter must be favored over antimatter
- C-symmetry must be broken  $X + Y \rightarrow \mathbf{B} > \overline{X} \rightarrow \overline{Y} + \overline{B}$
- CP-symmetry must be broken

$$(CP - X \rightarrow B_L + X \rightarrow B_R > Conjugates) \quad \overline{X} \rightarrow B_L + \overline{X} \rightarrow \overline{B}_R$$

• **Third Condition:** Excess **B** shouldn't be washed out – need this process to occur out of thermal equilibrium

![](_page_35_Picture_6.jpeg)

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# Electric dipole moments and CP violation

• Non-zero electric dipole moment in fundamental particle:  $H = -\frac{1}{2} \left( \vec{d} \cdot \vec{E} + \vec{\mu} \cdot \vec{B} \right)$  $\vec{d}$  $\vec{d}$ • P-violation, T-violation and therefore **CP**-violation Ù

 $\vec{\mu}$ 

## Electric dipole moments

![](_page_37_Figure_1.jpeg)

Chupp, Fierlinger, Ramsey-Musolf, and Singh, Rev. Mod. Phys. 91 1 (2019)

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## Neutron electric dipole moment

- Neutron has magnetic dipole moment: feels torque perpendicular to magnetic field
- Spin precesses with frequency  $h\nu = 2\mu_n B$
- If neutron had an EDM, spin precesses around electric field:

#### $h\nu = 2d_n E$

• Tactic: compare frequency with electric/ magnetic fields parallel/antiparallel

 $h\nu_{\pm} = 2\mu_n B \pm 2d_n E$   $\Delta\nu = \frac{4d_n E}{h}$ 

$$\Delta v = 7.5 \text{ nHz} \text{ for } d_n = 10^{-28} \text{ e-cm}, E = 75 \text{ kV/cm}$$

![](_page_38_Picture_8.jpeg)

## Limit on the neutron EDM

![](_page_39_Figure_1.jpeg)

- First measured by Smith, Purcell and Ramsey at ORNL
  - Studying parity violation in neutron scattering
- Neutron EDM extremely small: "Strong CP problem" in QCD Langrangian
- Motivates "axions" (dark matter candidate)

Smith, Purcell, and Ramsey, Phys Rev 108 120 (1957)

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## Ramsey's Method of Separated Oscillating Fields

- 1. Prepare "spin up" neutron
- 2. Apply 90° spin flip
- 3. Spin freely precesses by  $\omega \Delta t$
- 4. Apply 90° spin flip (to "spin down")
- 5. Measure "spin up" vs "spin down"

![](_page_40_Figure_6.jpeg)

![](_page_40_Figure_7.jpeg)

Т

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## Present best nEDM limit

![](_page_41_Picture_1.jpeg)

- Trapped 0.5 UCN/cc, stored for ~100 s
- $E \sim 10 \text{ kV/cm}$
- Result:  $d_n = -0.21 \pm 1.82 \times 10^{-26} \text{ e-cm}$

![](_page_41_Figure_5.jpeg)

## UCN from superfluid He (He-II)

- UCN production via "superthermal process"
- 8.9 Å (1 meV/11 K) cold neutron emits phonon in He-II, becomes UCN (500 Å/300 neV/3.5 mK)
- He-II cooled to ~1 K: Reverse process is suppressed by Boltzmann factor  $e^{-E/k_BT}$

![](_page_42_Figure_4.jpeg)

#### nEDM at the SNS

- Use helium to improve statistical precision
  - E: LHe = high voltage insulator, >85 kV/cm achieved
  - N: Higher densities using He-II
  - t: temp < 0.5 K allows storage times of  $\sim 1000$  s

![](_page_43_Figure_5.jpeg)

- and reduce systematic uncertainty
  - "Geometric phase": interaction of  $\vec{v} \times \vec{E} = \vec{B}_E$  and magnetic gradients
  - Limit leakage currents from electric field (which may produce magnetic fields)
  - Multiple measurement cells with opposite  $\vec{E} \cdot \vec{B}$
  - Multiple measurement techniques

Golub and Lamoreaux, Phys Rep 237 1 (1994) Ito et al, Rev Sci Instrum 87 045113 (2016)

#### SNS nEDM

- He-II is UCN source
- <sup>3</sup>He captures neutrons with strong spin-dependence
- He-II acts as a scintillator
- <sup>3</sup>He (95% polarized) used as a co-magnetometer

![](_page_44_Figure_5.jpeg)

![](_page_44_Picture_6.jpeg)

## Extracting the nEDM

- <sup>3</sup>He capture is strongly spindependent
  - 69 kb if spins anti-aligned
  - 11 b if spins aligned
- Time evolution of angle:

$$\theta_{n3} = |\gamma_n - \gamma_3| B_0 t \pm \frac{ed_n |E|}{\hbar} t$$

- $\cos(\omega_n \omega_3)$  variation in reaction rate
- Beat frequency: 9 Hz @ 30 mG
- <sup>3</sup>He as co-magnetometer
  - <sup>3</sup>He shielded by atomic electrons
  - <sup>3</sup>He precession measured directly by SQUID pickup

![](_page_45_Figure_11.jpeg)

## nEDM now being constructed

![](_page_46_Picture_1.jpeg)

## Outline

- P-violation and the weak interaction
- CP-violation and the electric dipole moment
- B-violation and oscillations

![](_page_48_Figure_0.jpeg)

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#### Neutron anti-neutron oscillations

$$\widehat{H} = \begin{pmatrix} m + \vec{\mu}(\vec{B} \cdot \vec{\sigma}) & \varepsilon \\ \varepsilon & m - \vec{\mu}(\vec{B} \cdot \vec{\sigma}) \end{pmatrix}$$

$$\Delta E = E_n - E_{\bar{n}} = 2\mu B$$
  
$$\Delta E (1 nT) \sim 10^{-22} \text{ MeV} \qquad \varepsilon < 10^{-29} \text{ MeV}$$

• Need "Quasi-free limit": Uncertainty principle =  $\Delta E \Delta t \ll \hbar$ 

2

• ILL: neutron TOF ~ 0.1 s, B~10 nT

• In this limit 
$$P_{n \to \overline{n}}(t) = \left(\frac{t_{free}}{\tau_{n \to \overline{n}}}\right)$$

#### Free $n \rightarrow \overline{n}$ Search at ILL

![](_page_50_Figure_1.jpeg)

Improving limits on  $n \rightarrow \overline{n}$ : European Spallation Source

- Similar time-averaged brightness to ILL
  - Pulsed = bkgd rejection
- >200 m beamguides possible
- Moderator, neutron optics, detector design study
  - ~1000x ILL possible

![](_page_51_Picture_6.jpeg)

Moderator configuration	Sensitivity Nt <sup>2</sup> in ILL units/yr
ESS TDR 2013	250
Option of large LD <sub>2</sub> source	550
Flat "pancake" with h = 3 cm	200
ANNI with BF1 Source	0.8

Adapted from M. Frost, INT 17-69W

## $\bar{n}$ annihilation detection

- Maintain "background free", improve efficiency
- Example "pion star"  $\overline{n} + {}^{12}C \rightarrow {}^{11}C$  $+3\pi^0 + \pi^+ + \pi^-$
- Goal background suppression  $< 10^{-8}$  Hz
- Key backgrounds:
  - Gamma backgrounds
    - Neutron capture
  - High energy n,p
  - Beta delayed n
  - Cosmic rays
    - Dominant, scales with size

![](_page_52_Figure_11.jpeg)

Detector study by A.R. Young, R.W. Pattie Jr, D.G. Phillips II

#### $n \rightarrow \overline{n}$ in Nuclei

![](_page_53_Figure_1.jpeg)

• Best limits from SuperKamiokande:  $\tau_{n \to \bar{n}} > 2.7 \times 10^8$  s

24 candidates, 24.1 expected backgrounds

Adapted from J. Barrow, BLV 2017 Abe et al, PRD 91 072006 (2015)

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## The Deep Underground Neutrino Experiment and Proposed $n - \bar{n}$ Search

![](_page_54_Picture_1.jpeg)

![](_page_54_Figure_2.jpeg)

Babu, Dev, Fortes, and Mohapatra- DOI:10.1103/PhysRevD.87.115019

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"Fundamental Symmetries" Exotic Beam Summer School, ORNL, July 24-29

#### Mirror neutron oscillations

- "Mirror matter": hidden sector dark matter candidate
- Multiple Baryon Numbers, with overall conservation of  $B_{tot} = B + B'$
- Experimental signature: oscillation into invisible "mirror neutrons"
  - All previous searches performed in ultracold neutron bottle experiments
- Various mechanisms to induce oscillations, using magnetic fields
  - some suggested to explain the neutron lifetime discrepancy
  - can consider mixing with antineutrons

$$\mathcal{H}_{int} = \begin{pmatrix} m + \mu \boldsymbol{\sigma} \cdot \boldsymbol{B} & \eta \boldsymbol{\sigma} \cdot [\boldsymbol{B} \pm \boldsymbol{B}'] + \varepsilon \\ \eta \boldsymbol{\sigma} \cdot [\boldsymbol{B} \pm \boldsymbol{B}'] + \varepsilon & m' + \mu' \boldsymbol{\sigma} \cdot \boldsymbol{B}' \end{pmatrix}$$

## Cold $n \rightarrow n'$ searches

![](_page_56_Figure_1.jpeg)

neutrons transform into mirror neutrons

- Add B field control P(n and high suppression "wall" to GP-SANS
- First small scale experiment this summer at SNS

![](_page_56_Picture_5.jpeg)

## Summary

- Tremendous opportunity to search for influence of new physics using low energy precision tests of fundamental symmetries
- Violation of CKM unitarity by 3  $\sigma$  demands new precision measurements using independent approaches
  - New precision neutron measurements on the horizon including Nab at SNS
- Matter asymmetry in universe remains one of the most important questions in science
  - Electric dipole moment searches such as nEDM at SNS will probe interesting parameter space for theories predicting CP-violation
  - Observation of neutron oscillations would demonstrate baryon number violating process
- Lots of exciting developments in neutrons and nuclei.... stay tuned!